

# The BLUE DOTS initiative

Elements for a roadmap

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for the Blue Dots team*



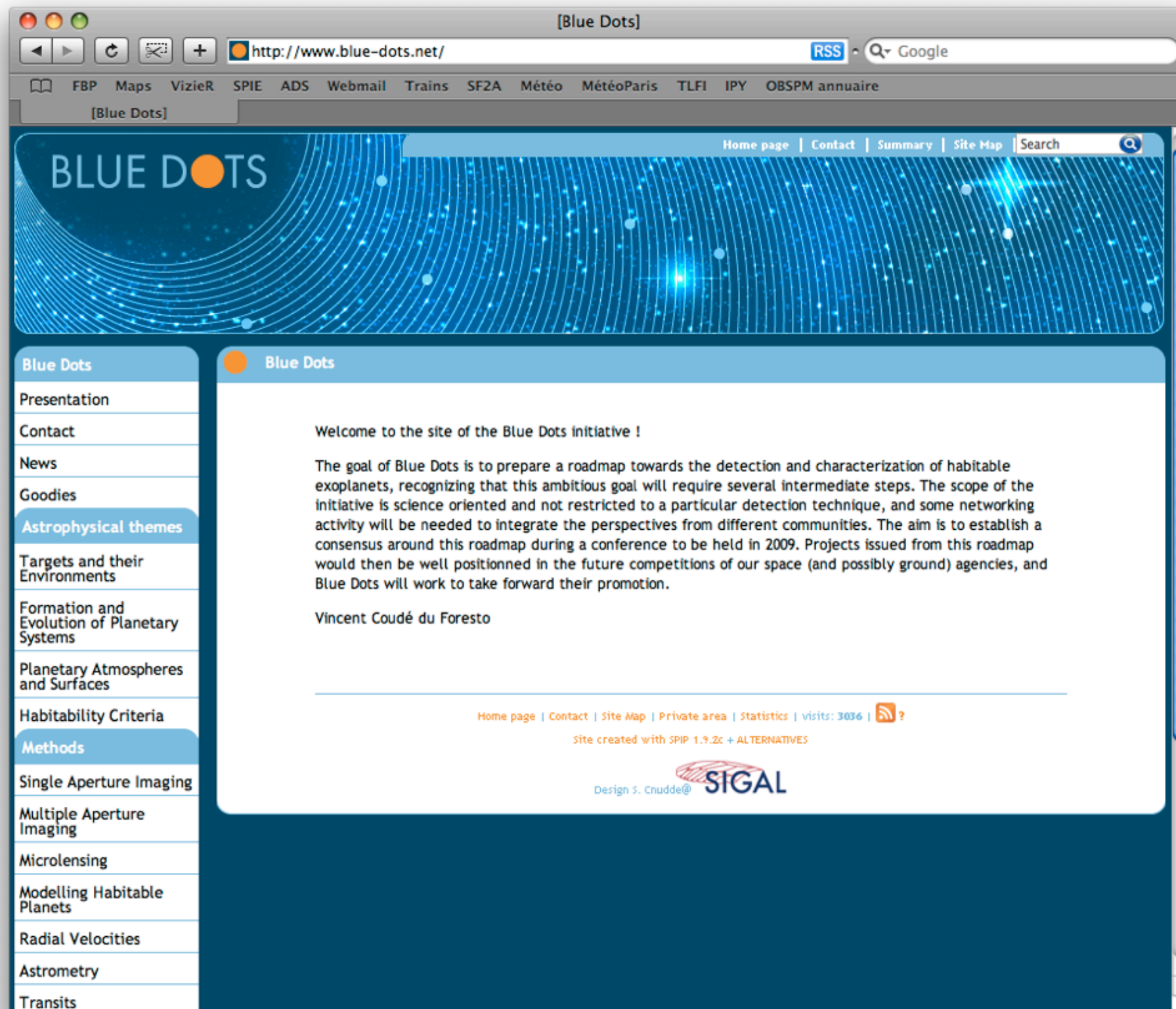
# Objectives of the initiative

- Contribute to building a community in Europe around the exoplanets theme
- Recognizing that the ultimate science goal (characterization of habitable exoplanet atmospheres) will require several intermediate steps...
- ...converge towards a strategy enabling a more coherent approach to calls for proposals in ground and space based projects...
- ...and most notably in time for the next CfP ESA linked to Cosmic Vision (expected in 2010)
- Beyond that: remain as a permanent structure to federate the community

# Deliverables Blue Dots Timeline

- Web site + intranet for information circulation
- Preliminary report to CNES february 2009
- Report v1 end March 2009
  - Call for feedback in the community
- Report stabilized by summer 2009
- Sep. 2009: Pathways conference (proceedings)
- 2010: structuration into European network (TBC)

# http://www.blue-dots.net



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# Methodology prospective

Setting a frame, grammar for roadmap elaboration

## *What do we want to know ?*

- For each astrophysical theme :
  - Establish prioritized list of key science questions
  - Physical facts needed to answer those key questions

## *What can we do ?*

- For each instrumentation theme:
  - Establish evolution of the quantity, quality of observables as a function of mission size

## *Then...*

- Compile this information in homogeneous tables
- The synthesis of these tables will help prepare the roadmap

# And also ...

- Tap into existing information
  - Expertise within WGs
  - Input from existing white papers
  - Proactive external collection when needed
  - Spontaneous external contributions always possible
- Clearly separate matters of consensus from matters of debate
- Deliberately introduce specific missions at a later stage of the prospective

# What do we want to know ?

- Habitability criteria
- Planetary atmospheres & surfaces
- Formation & evolution of planetary systems
- Targets & their environments



# What do we want to know ?

- Habitability criteria
- Planetary atmospheres & surfaces
- Formation & evolution of planetary systems
- Targets & their environments
  - *What are the physical properties (including mass and age) of the target stars?*
  - *What are the radiative properties (light and particles,  $T_{\text{eff}}$  ,  $L_{\text{bol}}$ ) of the target stars?*
  - *What is the time-variation of such emissions ?*
  - *What are the characteristics of the stellar immediate surroundings (i.e., zodiacal dust, companion stars, brown dwarfs or giant planets)?*
  - *What are the stellar properties (mass, chemical composition) influencing the existence of telluric planets?*

# What can we do

- A step by step approach:
  - 1 – Statistical study of planetary objects
  - 2 – Designate sources suitable for spectroscopic follow-up
  - 3 – Carry out spectroscopic characterization

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*These define different science potential levels  
which can be achieved on different object classes =>  
different difficulties*

# Methods and scales

- Methods:
  - RV: Radial Velocities
  - μlensing
  - Transit photometry
  - Single Aperture Imaging
  - Multiple Aperture Imaging
  - Astrometry
  - *Modelling planetary structures & atmospheres*

# Methods and scales

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  - Astrometry
  - *Modelling planetary structures & atmospheres*
- Scales:
  - E (existing)
  - G (30M€, 5 years)
  - M (450M€, 10 years)
  - L (650M€, 15 years)
  - XL (> 1G€, > 20 years)

# Science potential level

## Methods and scales

	1	2	3	4	5
	Giant (close / young)	Giants (others)	Telluric (others)	Telluric HZ (M)	Telluric HZ (others)
Astrometry	2	2	2		2
MAI	3		3	3	3
SAI	3	3		3	3
Transit	3	2 ?	3 ?	3	1
μlensing		1	1	1	1
RV	2	2	2	2	2

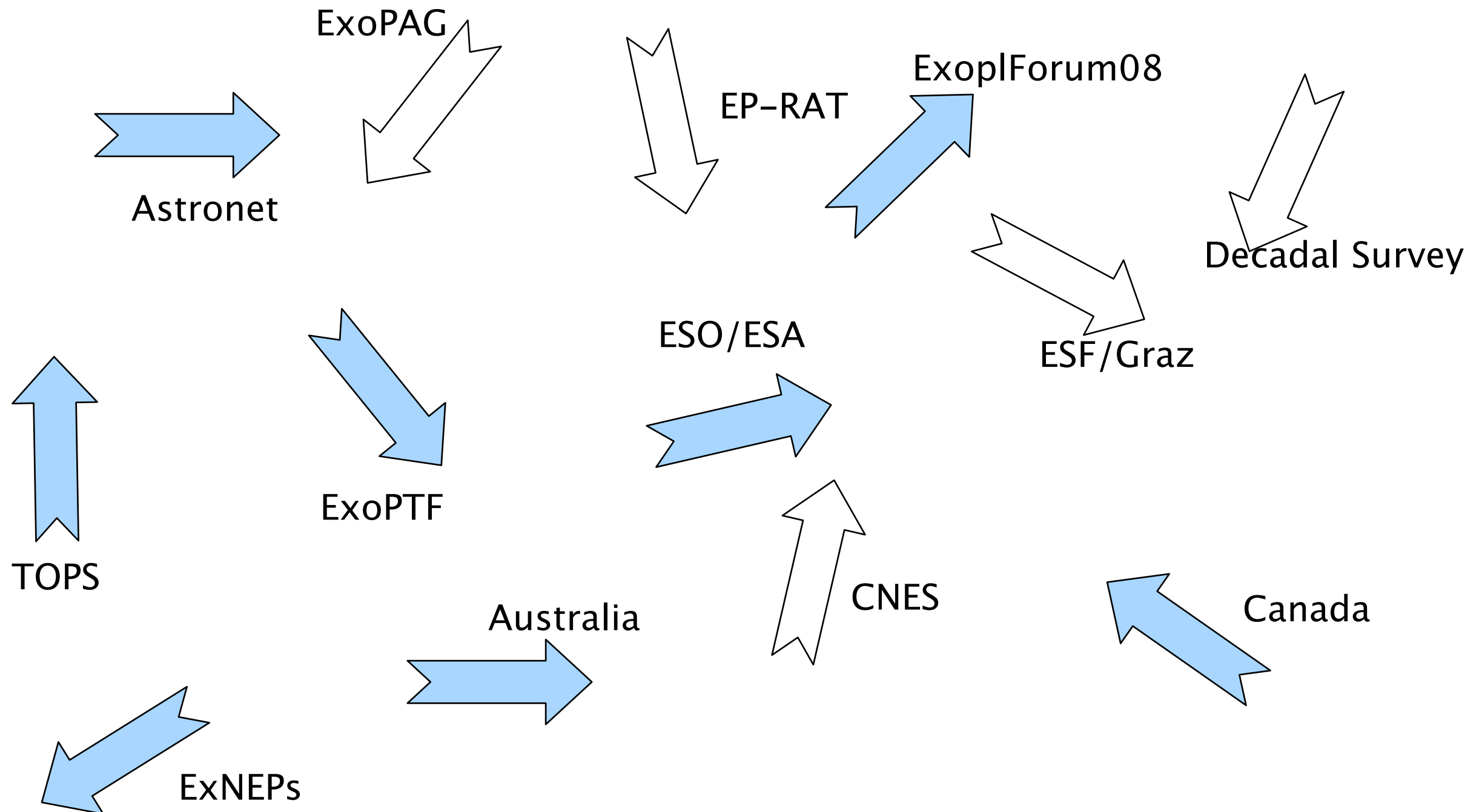
- Cornerstone questions:
  - Can telluric habitable planets be identified from the ground by RV ?
  - Should we search for habitable planets around M stars ?
  - Is spectroscopic characterization of the atmosphere of telluric exoplanets possible by transit spectroscopy ?
  - Do we need to solve the exozodi issue ? If yes, how best to solve it ?

# Timeline methods

	SPL	Existing	G	M	L - XL
Astrometry	2		Giants (far)		Telluric
MAI	3		Hot giants		All others
SAI	3	Young giants	Giants (far)		Telluric
Transit	2 - 3 ?	Close giants		All others	
μlensing	1	Giants		Habitable telluric	
RV	2	Giants, telluric	Habitable telluric		

- Cornerstone questions:
  - Do we need precursors to large flagship missions ?
  - How relevant will be, when it flies, a large flagship mission defined now ?
  - Are institutional structures compatible with an ambitious exoplanetary program ?

# Exoplanet roadmaps...



*Courtesy J. Schneider*



<http://www.pathways2009.net>



# Opportunity for satellite meetings

- Benefit from the full exposure and logistics of the Pathways conference
- Extra sessions in late afternoon
- Deadline for proposals May 20th
- See <http://www.pathways2009.net/satellite.html> for details

